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Putting the D in MRMHD a prescription for all that ails ideal MHD!

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Why and how to fix Ideal MHD?

- Ideal MHD is overconstrained
 - No heat transport along field lines
 - No reconnection so islands or chaos cannot form
 - Thus inapplicable to hot and 3D plasmas!
- Fix by removing the bad constraints and keeping the good, doing more with less!





MRxMHD: M stands for **Multi-region** (aka waterbag) Rx stands for **Relaxed**; ...D stands for **Dynamics**





Fundamental postulates of new general reformulation of MHD:

□ ∃ transport interfaces, \mathcal{I}_i or $\Gamma_{i,j}$, or $\partial \Omega_{i,j}$ (e.g. nested tori or island separatrices), that act like sheets of ideal-MHD plasma

□ Plasma relaxes (in some generalized Taylor sense) in regions \mathcal{P}_i (or Ω_i) bounded by the interfaces

Only a subset of ideal-MHD invariants apply



SPEC (currently) uses MRxMHS, not MRxMHD:

MRxMHS = Multi-region Relaxed MagnetohydroStatics (i.e. equilibrium theory)

- Taylor relaxation *energy* principle
- constant pressure in each region

MRxMHD = Multi-region Relaxed Magnetohydro*Dynamics*

New approach:: ^{*} use *Hamilton's Principle* — stationarity of time-integrated *Lagrangian*

⇒ constant *temperature* in each region

supports sound waves within relaxation regions as well as radially compressible and Alfvén modes + *tearing*

can treat *development* of resonant current sheets

can add equilibrium flow to SPEC and will be basis for a new time-evolution waterbag code

Ref. Stuart Hudson's talk yesterday



MRxMHD Lagrangian is *kinetic energy* minus MHD *potential energy* + constraint terms:

• MHD Lagrangian density in region *i*

$$\mathcal{L}^{\rm MHD} = \rho \frac{v^2}{2} - \frac{p}{\gamma - 1} - \frac{B^2}{2\mu_0}$$

• Constrained Lagrangian in region *i*

$$L_i = \int_{\Omega_i} \mathcal{L}^{\text{MHD}} dV + \tau_i (S_i - S_{i0}) + \mu_i (K_i - K_{i0})$$

• Helicity and entropy macroscopic invariants

$$K_i \equiv \int_{\Omega_i} \frac{\mathbf{A} \cdot \mathbf{B}}{2\mu_0} \, dV \qquad S_i \equiv \int_{\Omega_i} \frac{\rho}{\gamma - 1} \ln\left(\kappa \frac{p}{\rho^{\gamma}}\right) dV$$



In varying action, ρ is constrained holonomically to the displacement $\boldsymbol{\xi}$ of each fluid element:

• Mass conserved *microscopically*, i.e. pointwise

$$\delta \rho = -\nabla \cdot (\rho \boldsymbol{\xi}) \text{ in } \Omega_i$$

- Helicity and entropy constrained *macroscopically*, throughout Ω_i, using Lagrange multipliers μ_i and τ_i, while *p* and **A** are free fields
- Including vacuum field energy, total Lagrangian is

$$L = \sum_{i} L_{i} - \int_{\Omega_{v}} \frac{\mathbf{B} \cdot \mathbf{B}}{2\mu_{0}} \, dV$$

• Setting variation of action to 0 gives EL equations: $\delta \int Ldt = 0$



Equations within Ω_i

• Mass conservation (microscopic constraint)

$$\frac{\partial \rho}{\partial t} = -\boldsymbol{\nabla} \boldsymbol{\cdot} (\rho \mathbf{v})$$

• $\delta p \Rightarrow$ Isothermal equation of state

$$p = \tau_i \rho$$
 (N.B. $\tau_i = C_{\mathrm{s}i}^2$)

δA ▷ Beltrami equation
∇×B = μ_iB (N.B. ⇒ j×B = 0)
ξ ▷ Momentum equation (Euler fluid)

$$\rho\left(\frac{\partial \mathbf{v}}{\partial t} + \mathbf{v} \cdot \nabla \mathbf{v}\right) = -\nabla p$$



Equations on interface $\Gamma_{i,j}$

• **ξ** ♀ Force balance

$$\left[p + \frac{B^2}{2\mu_0}\right]_{i,j} = 0$$

Surface constraints

 $\mathbf{n}_i \cdot \mathbf{B} = 0 \quad \text{on } \partial \Omega_i$ $\mathbf{n}_i \cdot [\![\mathbf{v}]\!]_{i,j} = 0 \quad \text{on } \partial \Omega_{i,j}$

• Complete set of equations, consistent because derived from single scalar function *L*



Proving the MRxMHD pudding:

- Q1) What is the MRxMHD spectrum and what are the effects of field-line curvature and <u>equilibrium mass flow</u> on stability?
- Q2) When are the <u>current sheets</u> topologically stable towards internal plasmoid formation (reconnection)?
- Q3) When do unstable modes saturate at a low level or develop nonlinearly into explosive events?



What happens in static or adiabatic limit?

•
$$\partial_t \to \mathbf{0} \quad \nabla \cdot (\rho \mathbf{v}) = 0 \quad \mathbf{v} \cdot \nabla \mathbf{v} = -\tau_i \nabla \ln \rho$$

only solutions valid for any flowline configuration, from nested surfaces to arbitrarily chaotic, are

$$\rho = \rho_{0i} \exp\left(-\frac{v^2}{2\tau_i}\right) \diamondsuit p = p_{0i} \exp\left(-\frac{v^2}{2\tau_i}\right)$$

(N.B. *incompressible* in limit $v/C_s \rightarrow 0$) and

$$\boldsymbol{\nabla} \times \mathbf{v} = \alpha_{0i} \exp\left(-\frac{v^2}{2\tau_i}\right) \mathbf{v}$$

Almost isomorphous to **B** equation: should be implementable in SPEC. Derivable variationally — N. Sato

Switch on slab boundary ripple to study Resonant Magnetic Perturbations (RMPs)

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MRxMHD Hahm-Kulsrud-Taylor (HKT): Rippled Slab Model for resonant current sheets

Australian National University 2-region MRXMHD model*



*From APS DPP 2014 poster,

with Finn-Antonsen helicity

• Simple slab model for resonant current sheet formation near x = 0 in response to symmetrical periodic perturbation at boundaries $x = \pm a$

Hahm & Kulsrud (HK), Phys. Fluids '85, found 2 solutions:

• shielding current sheet on x = 0 (shown in red)

$$\psi = aB_y^a \left[\frac{x^2}{2a^2} + \frac{\alpha}{\sinh(ka)} |\sinh(kx)| \cos(ky) \right]$$

• island with no current sheet

$$\psi = aB_y^a \left[\frac{x^2}{2a^2} + \frac{\alpha}{\cosh(ka)}\cosh(kx)\cos(ky)\right]$$

where B^{a}_{y} is [unperturbed poloidal field] at boundaries and $\alpha \ll 1$

A good test case for MRxMHD:

- Linearity of Beltrami equation leads to easily solvable, linear GS equation (Poisson in small-μ limit.)
- Symmetry about, and straightness of, current sheet at x = 0: gives most
- > geometrically simple 2-region relaxation scenario:
- Switch-on: *ripple* on upper and lower boundaries slowly increased from zero (plane slab) to final amplitude
- A *shielding current* sheet at *x* = 0 resonance develops
- Kruskal-Kulsrud damping: evolution through equilibria
- Connect equilibrium sequence by helicity conservation

Grad-Shafranov equation for force-free field in slab geometry: $\mathbf{B} = \nabla z \times \nabla \psi + F(\psi) \nabla z \qquad \nabla^2 \psi + FF' = 0$ $\nabla \times \mathbf{B} = \mu \mathbf{B}$ (Beltrami equation) is satisfied by requiring: $abla^2\psi=\mu F$ with $F(\psi)=C-\mu\psi$, giving $(abla^2+\mu^2)\psi=C$ General Solution: $\psi = \overline{\psi} + \frac{\overline{F}}{B_0}\psi_0(x|\mu) + \widehat{\psi}(x,y)$ where $\overline{\psi}$ is cross-sectional average of ψ , $\psi_0(x|\mu) \equiv \frac{B_0}{-1}(1-\cos\mu x)$ is plane slab solution, \overline{F} is the cross-sectional average of B_z , and $\widehat{\psi}$ obeys a *homogeneous* Beltrami equation $(\nabla^2 + \mu^2)\widehat{\psi} = 0$ with boundary conditions such that ψ is constant on boundary and on cuts.

Grad-Shafranov-Beltrami equations

National University Extension of HK shielding solution Helicity conservation requires three extensions of HK solution: Instead of the HK harmonic component ψ_1 we use ansatz

$$\widehat{\psi}(x,y) \equiv \frac{2\alpha\psi_a}{\sinh k_1 a} \left(|\sinh k_1 x| \cos ky + \gamma_{\rm S} \frac{k_1}{\mu} |\sin \mu x| \right) - \overline{\psi} \cos \mu x$$

where:

1. $\hat{\psi}$ is a solution of the *Beltrami equation* $(\nabla^2 + \mu^2)\hat{\psi} = 0$ It is only *harmonic* in the *small-µ limit*. Likewise

$$k_1(\mu) \equiv (k^2 - \mu^2)^{1/2} \rightarrow k \text{ only as } \mu \rightarrow 0$$

- 2. The term in γ_s was introduced in Dewar *et al*. 2013 to allow control of the *total current* in the sheet
- 3. The term in $\overline{\psi}$ is required for poloidal flux conservation



- In plane slab, *before* ripple is turned on, the *unperturbed* equilibrium flux function is $\psi_0(x|\mu_0) \equiv \frac{B_0}{\mu_0}(1 - \cos \mu_0 x)$
 - As amplitude parameter α is increased from 0, μ must *change* to preserve helicity and fluxes:





• HK implicitly assumed the total current in the sheet was zero, but MRxMHD switch-on shows there is a nonzero total current $J = \frac{2\alpha\psi_a k_1\lambda}{\sinh k_1a}\gamma_{\rm S}$ proportional to $\gamma_{\rm S}$:





Current sheet reverses for small perturbations



Fully shielded case: Plots of the jump in the gradient of ψ , vs. *y* for $\mu_0 = 1.4$ and selected small values of α , showing the occurrence of current-density reversal for the two smallest values.



Current reversals cause 1/2 islands!



y = 1.0 - 0.5 0.0 - 0.5 1.0 = $5.0 \times 10^{-0.5}$ reversal 2-region MRxMHD Hahm-Kulsrud model: mirror-image ripple top and bottom excites modulated current sheet at x = 0





-luxes and transform

2

0.000

-0.0006

0.0004





Poloidal flux as a function of x_0 (= xalong y-axis), showing discontinuity in slope transform at x = 0 caused by 0.0006current sheet 0.0004

Toroidal flux as a function of xalong y-axis, showing discontinuity at x = 0 caused by half-island.

> Rotational transform (1/q) $_{10^{5}x_{0}} \quad \Psi'(x_{0})/\Phi'(x_{0})$ showing jump or large slope near $x_{0} = 0$.



Fluxes and transform II



α = **0.005** (Dashed curves

are for plane slab, $\alpha = 0$)

Discontinuity in toroidal flux has gone as there are no half-islands above a threshold in α c. 0.0045

 Φ, Φ_0

2

4

 $10^5 x_0$

0.00005

.00005



⁹ Much stronger jump in rotational transform



Full *t*-dependence: linear modes in slab



Alexis Tuen's MSc thesis 2016



First two eigenvalues, + incompressible approximation at very small $\lambda \equiv \omega^2$



- Dewar, Tuen, Hole: Plasma Phys. Control. Fusion 59, 044009, (2017)
- Growth rate zero if wall or **k.B**=0 is at interface



- Multi-region generalization of Taylor relaxation has been extended to a self-consistent dynamics through Hamilton's Principle of Stationary Action.
- A rippled slab model has been used to illustrate the formation of a resonant current sheet as boundary ripple is switched on
- For very small ripple amplitudes current reversal occurs in the current sheet and unperturbed sheared magnetic field exhibits topological change, with small half-islands, locking rotational transform to resonant value
- For larger ripple amplitude the rotational transform jumps across the current sheet



General Conclusion

- Action-based MRxMHD shows great promise
 - Very simple
 - Includes reconnection and flow in natural way
- 1st: check physical reasonability of predictions in simple models
- 2^{nd:} Extend SPEC to flow; build new timeevolution and normal mode codes